

# Final Examination Paper for Electrodynamics-I

Date: Friday, Nov 02, 2007,

Time: 09:00 - 15:00

[Solutions]

Allowed help material: *Physics and Mathematics handbooks or equivalent*

Note: Please explain your reasoning and calculations clearly

Questions:	1	2	3	4	5	6	Total
Marks:	13	13	14	13	13	14	80

1. (a) Consider an electric field  $\vec{E} = \hat{i}x + \hat{j}z + \hat{k}(f(x, y) + z^2)$ . Determine  $f(x, y)$  and compute the total charge contained in a cube specified by  $0 \leq x, y, z \leq l$ .
- (b) The electrostatic potential of a neutral atom can be modelled by

$$\Phi(\vec{r}) = \frac{q}{r} e^{-r/a}$$

where  $q = Ze$  is the atomic charge. Find the charge distribution  $\rho$  that produces this potential and show that the total charge is zero. (You may use  $\vec{\nabla}r = \hat{r}$  and  $\nabla^2 r = 2/r$ )

**Solution** (points: 6+7)

a) This is an electrostatic field with  $E_x = x$ ,  $E_y = z$ ,  $E_z = f(x, y) + z^2$  and should satisfy  $\vec{\nabla} \times \vec{E} = 0$ . In terms of components of  $\vec{E}$  this gives  $\partial E_i / \partial x^j - \partial E_j / \partial x^i = 0$  for the indices  $i$  and  $j$  taking the values  $x, y, z$ , which, in turn, leads to  $\partial f / \partial x = 0$  and  $\partial f / \partial y - 1 = 0$ . The unknown function  $f(x, y)$  is therefore given by  $f = y + c$  for an arbitrary constant  $c$ . So we have,  $\vec{E} = \hat{i}x + \hat{j}z + \hat{k}(y + z^2 + c)$ . From this, we can compute the charge density using  $\vec{\nabla} \cdot \vec{E} = 4\pi\rho$  and get  $\rho = (2z + 1)/4\pi$ . The total charge is then given by

$$Q = \frac{1}{4\pi} \int_0^l dx \int_0^l dy \int_0^l dz (1 + 2z) = \frac{1}{4\pi} (l^3 + l^4)$$

(the total charge can also be computed using the Gauss law)

b) The charge distribution can be determined using  $\nabla^2 \Phi = -4\pi\rho$  and is given by

$$\rho = qe^{-r/a} \left( \delta(r) - \frac{1}{4\pi} \frac{1}{ra^2} \right)$$

This clearly corresponds to a positive nuclear point charge and a negative electronic charge cloud surrounding it. The total charge is given by

$$Q = \int d^3x \rho = q - \frac{q}{a^2} \int_0^\infty e^{-r/a} r dr = 0$$

where we have used  $\int d^3x = \int \sin\theta d\theta \int d\phi \int r^2 dr = 4\pi \int r^2 dr$  because of the spherical symmetry of the problem, along with  $\int_0^\infty e^{-r/a} dr = a$ , which on differentiating with respect to "a" gives  $\int_0^\infty e^{-r/a} r dr = a^2$ .

2. Consider the boundary between two media of dielectric constants  $\epsilon_1$  and  $\epsilon_2$  and let the electric displacement vectors on the two sides of the boundary be denoted by  $\vec{D}_1$  and  $\vec{D}_2$ , and the polarization densities by  $\vec{P}_1$  and  $\vec{P}_2$ , respectively. In the absence of free charges on the boundary, Maxwell equations are  $\vec{\nabla} \cdot \vec{D} = 0$  and  $\vec{\nabla} \times \vec{E} = 0$ .

- (a) Use these equations to investigate the continuity of the normal and tangential components of  $\vec{D}$  and  $\vec{E}$  across the boundary.
- (b) Show that the polarization surface charge density that develops on the boundary is given by

$$\sigma_{pol} = (\vec{P}_1 - P_2) \cdot \hat{n},$$

where  $\hat{n}$  is a unit normal to the boundary.

**Solution** (points: 8+5)

a) In general, in steady state the fields  $\vec{D}$  and  $\vec{E}$  satisfy  $\vec{\nabla} \cdot \vec{D} = 4\pi\rho_f$  and  $\vec{\nabla} \times \vec{E} = 0$ , where  $\rho_f$  is the density of free charges. To explore the behaviour of the normal component of  $\vec{D}$ , first, draw a small, so called, ‘‘Gaussian pill-box’’ of height  $h$  across the boundary. The top and bottom faces of the pill-box have areas  $\Delta S$  each and are parallel to the boundary surface. Denote the value of the displacement field on the bottom face of the box by  $\vec{D}_1$  and on the top face of the box by  $\vec{D}_2$ . The unit normals to these faces are  $\hat{n}_1$  and  $\hat{n}_2$  ( $\hat{n}_2 = -\hat{n}_1 = \hat{n}$ ). Then, integrating  $\vec{\nabla} \cdot \vec{D}$  over the pill-box volume and using the divergence theorem gives,

$$\lim_{h \rightarrow 0} \int_{pill\text{-}box} d^3x \vec{\nabla} \cdot \vec{D} = \lim_{h \rightarrow 0} \int_{\partial(pill\text{-}box)} \vec{d}s \cdot \vec{D} = (\vec{D}_2 \cdot \hat{n}_2 + \vec{D}_1 \cdot \hat{n}_1) \Delta S = 4\pi\sigma_f \Delta S$$

where  $\sigma_f$  is the density of free charges on the boundary and the contribution from the sides have dropped in the limit  $h \rightarrow 0$ . Hence we have,

$$(\vec{D}_2 - \vec{D}_1) \cdot \hat{n} = 4\pi\sigma_f$$

In our problem, there are no free charges on the boundary and hence  $(\vec{D}_2 - \vec{D}_1) \cdot \hat{n} = 0$ . So the normal component of the  $\vec{D}$  field is continuous across the boundary.

To investigate the behaviour of the tangential component of the field, let us now replace the pill-box by a rectangular loop that has its longer sides of length  $\Delta l$  parallel to the surface and its shorter sides of height  $h$  perpendicular to the surface and going through it. A unit vector along the lower side of the rectangle is  $\hat{t}_1$  and one along the upper side is  $\hat{t}_2$ , both being parallel to the surface. Picking an orientation along the loop, one has  $\hat{t}_2 = -\hat{t}_1 = \hat{t}$ . Integrate  $\vec{\nabla} \times \vec{E}$  over the loop area to get

$$\lim_{h \rightarrow 0} \int_{loop\text{ area}} \vec{\nabla} \times \vec{E} \cdot \vec{d}s = \lim_{h \rightarrow 0} \int_{loop} \vec{E} \cdot \vec{dl} = (\vec{E}_2 \cdot \hat{t}_2 + \vec{E}_1 \cdot \hat{t}_1) \Delta l$$

where in the limit  $h \rightarrow 0$  we have dropped the contributions from the sides of the loop. This is true for all orientations of the loop, or equivalently, for all unit tangent vectors  $\hat{t}$  to the surface. Therefore we have,  $(\vec{E}_2 - \vec{E}_1) \cdot \hat{t} = 0$  or equivalently,

$(\vec{E}_2 - \vec{E}_1) \times \hat{n} = 0$ . Hence the tangential component of  $\vec{E}$  is continuous across the surface.

b) Applying the pill-box construction above to  $\vec{\nabla} \cdot \vec{E} = 4\pi\rho$  where  $\rho = \rho_f + \rho_{nf}$  includes both free and non-free (i.e., bound) charges, one gets,

$$(\vec{E}_2 - \vec{E}_1) \cdot \hat{n} = 4\pi(\sigma_f + \sigma_{nf})$$

Now, for  $\sigma_f = 0$  on the boundary,  $\vec{D} \cdot \hat{n}$  is continuous across the boundary. Using this and the relation  $\vec{D} = \vec{E} + 4\pi\vec{P}$ , one has,

$$(\vec{E}_2 - \vec{E}_1) \cdot \hat{n} = 4\pi(\vec{P}_1 - \vec{P}_2) \cdot \hat{n}$$

From this we can read off the surface density of non-free charges on the boundary which are due to the polarization of the media as  $\sigma_{nf} = \sigma_{pol} = (\vec{P}_1 - \vec{P}_2) \cdot \hat{n}$ . In short, this directly follows from the fact that  $\vec{D}$  across the surface is continuous while the discontinuity in  $\vec{E} \cdot \hat{n}$  is given by the surface charge density and  $\vec{D} = \vec{E} + 4\pi\vec{P}$ .

3. (a) Using the expansion

$$\frac{1}{|\vec{x} - \vec{x}'|} = \sum_{l=0}^{\infty} \sum_{m=-l}^l \frac{4\pi}{2l+1} \frac{r'^l}{r^{l+1}} Y_{lm}^*(\theta', \phi') Y_{lm}(\theta, \phi)$$

develop the multipole expansion of the potential  $\Phi(\vec{x})$  due to a localized charge distribution  $\rho(\vec{x}')$  in terms of the multipole moments  $q_{lm}$  of  $\rho$ . Discuss how and under what conditions this expansion can be used to simplify a problem.

- (b) Show that for a spherically symmetric charge distribution, all multipole moments beyond the monopole moment vanish.
- (c) Show that if the charge distribution has axial symmetry (that is, it is invariant under rotations about the z-axis), then the only non-zero multipole moments are  $q_{l0}$ .
- (d) Using the above results, for two point charges  $q$  and  $-q$  placed on the z-axis at  $z = a$  and  $z = -a$ , compute the non-vanishing component of the dipole moment (given  $Y_{10} = (\sqrt{3/4\pi}) \cos \theta$ ).

**Solution** (points: 5+3+3+3)

a) The potential due to a localized charge distribution is given by

$$\Phi(\vec{x}) = \int d^3x' \frac{\rho(\vec{x}')}{|\vec{x} - \vec{x}'|}$$

Using the expansion given in the question, it becomes,

$$\Phi(\vec{x}) = \sum_{l=0}^{\infty} \sum_{m=-l}^{m=l} \frac{4\pi}{2l+1} q_{lm} \frac{Y_{lm}(\theta, \phi)}{r^{l+1}}$$

This is the multipole expansion of the potential in terms of the multipole moments  $q_{lm}$  of the charge distribution given by

$$q_{lm} = \int d^3x' \rho(\vec{x}') r'^l Y_{lm}^*(\theta', \phi')$$

The multipole expansion allows us to parametrize the charge distribution in terms of its multipole moments. Further, the contribution of a moment  $q_{lm}$  to the potential falls off as  $1/r^{l+1}$ . Therefore, at large distances from a localized charge distribution, only a few non-zero multipole moments with the lowest values of  $l$  make significant contributions to  $\Phi$  and are relevant. The remaining moments could be neglected. This allows us to parametrize even complicated charge distributions in terms of a few lowest  $l$  multipole moments. The condition under which this approximation is valid is that the distance to the observation point (at which  $\Phi$  is measured) is much larger as compared to the size of the charge distribution.

b) For a spherically symmetric charge distribution,  $\rho(\vec{x}) \equiv \rho(r, \theta, \phi) = \rho(r)$ , independent of the angular variables. Therefore we can write the multipole moments as a product of two integrals,

$$q_{lm} = \int_0^\infty r'^2 dr' \rho(r') r'^l \int_0^\pi \sin \theta' d\theta' \int_0^{2\pi} d\phi' Y_{lm}^*(\theta', \phi')$$

Since  $Y_{00}(\theta', \phi') = 1/\sqrt{4\pi}$ , we can insert  $Y_{00}(\theta', \phi')\sqrt{4\pi} = 1$  in the angular integration. Now, from the orthogonality property of spherical harmonics it follows that the angular integral is proportional to  $\delta_{l0}$  and hence vanishes for all  $l \geq 1$  ( $l = 0$  being the monopole moment).

c) In the case of axial symmetry about the  $z$ -axis,  $\rho$  is independent of the azimuthal coordinate  $\phi$ . In this case,

$$q_{lm} = (\text{const}) \int_0^\infty r'^2 dr' \int_0^\pi \sin \theta' d\theta' \rho(r', \theta') r'^l P_l^m(\cos \theta') \int_0^{2\pi} d\phi' e^{-im\phi'}$$

where we have used the fact that  $Y_{lm}(\theta, \phi) = (\text{const})P_l^m(\cos \theta)e^{im\phi}$ . Now, the  $\phi$  integral gives a  $\delta_{m0}$ . Thus the only non-vanishing moments in this case are  $q_{l0}$ .

d) In this case the charge density is given by  $\rho = q\delta(x')\delta(y')(\delta(z' - a) - \delta(z' + a))$ . The three components of the dipole moment are  $q_{1m}$ , for  $m = 1, 0, -1$ . Since the problem has axial symmetry about the  $z$ -axis, the only non-vanishing component is  $q_{10}$  which is now given by (using the expressions for  $Y_{10}$ ,  $\rho$  and noting that  $r' = \sqrt{x'^2 + y'^2 + z'^2}$ )

$$\begin{aligned} q_{10} &= \int d^3x' \rho(\vec{x}') r' Y_{10}^*(\theta', \phi') = \sqrt{\frac{3}{4\pi}} \int dx' \int dy' \int dz' \\ &\quad \times q\delta(x')\delta(y')(\delta(z' - a) - \delta(z' + a)) \sqrt{x'^2 + y'^2 + z'^2} \cos \theta \\ &= \sqrt{\frac{3}{4\pi}} \int dz' q(\delta(z' - a) - \delta(z' + a)) |z'| = aq\sqrt{\frac{3}{\pi}} \end{aligned}$$

Note that  $\delta(x')$  and  $\delta(y')$  force  $\vec{x}'$  to be in the  $z$  direction and hence,  $\theta = 0$ .

4. (a) Discuss the consistency of the magnetostatic equation  $\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{J}$  with the continuity equation  $\vec{\nabla} \cdot \vec{J} = -\frac{\partial \rho}{\partial t}$ .
- (b) Show that the work done by a magnetic field  $\vec{B}$  on a charged particle, moving with velocity  $\vec{v}$  under the influence of  $\vec{B}$ , is zero.
- (c) Starting with the magnetostatic equation given in part (a) derive Ampere's law for a straight conducting wire carrying current  $I$ .

**Solution** (points: 5+4+4)

a) The magnetostatic equation implies that  $\vec{\nabla} \cdot \vec{J} = \frac{c}{4\pi} \vec{\nabla} \cdot (\vec{\nabla} \times \vec{B}) = 0$ . Hence, it is consistent with the continuity equation only when  $\frac{\partial \rho}{\partial t} = 0$ . On the other hand, from the Gauss law equation it follows that  $\frac{\partial \rho}{\partial t} = \frac{1}{4\pi} \vec{\nabla} \cdot \left(\frac{\partial \vec{E}}{\partial t}\right)$ . Therefore, if the magnetostatic equation is modified to  $\vec{\nabla} \times \vec{B} - \frac{1}{c} \left(\frac{\partial \vec{E}}{\partial t}\right) = \frac{4\pi}{c} \vec{J}$ , it becomes fully consistent with the continuity equation.

b) The elemental work done on a charged particle of velocity  $\vec{v}$  moving in a magnetic field is  $dW = \vec{F} \cdot d\vec{x} = \frac{q}{c} (\vec{v} \times \vec{B}) \cdot d\vec{x}$ . But in this case,  $d\vec{x} = \vec{v} dt$  and since  $\vec{v} \times \vec{B}$  is perpendicular to  $\vec{v}$  (and of course also to  $\vec{B}$ ), it follows that  $(\vec{v} \times \vec{B}) \cdot \vec{v} = 0$ .

c) On a plane perpendicular to the current carrying conductor consider a disc of radius  $r$  centred at the conductor. Integrate the magnetostatic equation over the area of this disc. Then  $\int d\vec{S} \cdot \vec{J} = I$  and  $\int d\vec{S} \cdot (\vec{\nabla} \times \vec{B}) = \oint \vec{B} \cdot d\vec{l} = 2\pi r B_\phi$ , where,  $B_\phi$  is the component of  $\vec{B}$  tangent to the boundary of the disc. From the symmetry of the problem, this clearly is the only non-vanishing component of  $\vec{B}$ . Hence,  $\vec{B} = (2I/rc)\hat{\phi}$ .

5. (a) Starting from Maxwell equations, derive the wave equation satisfied by the vector potential  $\vec{A}$  in the Lorenz gauge. (You may need the vector identity  $\vec{\nabla} \times (\vec{\nabla} \times \vec{A}) = -\nabla^2 \vec{A} + \vec{\nabla}(\vec{\nabla} \cdot \vec{A})$ )
- (b) The equation for the vector potential  $\vec{A}$  in the Lorenz gauge and in the presence of a current source has a solution

$$\vec{A}(\vec{x}, t) = \frac{1}{c} \int d^3x' \frac{[\vec{J}(\vec{x}', t')]_{ret}}{|\vec{x} - \vec{x}'|}$$

in terms of the retarded time  $t' = t - |\vec{x} - \vec{x}'|/c$ . For a sinusoidal source term,  $\vec{J}(\vec{x}, t) = \vec{J}(\vec{x})e^{-i\omega t}$ , write down and discuss the nature of the solution in the "near zone" and the "far zone" approximations.

**Solution** (points: 6+7)

a) Start with the Maxwell equation containing the source term  $\vec{J}$  and substitute for the electric and magnetic fields in terms of the potentials,  $\vec{E} = -\vec{\nabla}\Phi - \frac{1}{c}(\partial\vec{A}/\partial t)$  and  $\vec{B} = \vec{\nabla} \times \vec{A}$ . This gives

$$\nabla^2 \vec{A} - \frac{1}{c^2} \frac{\partial^2 \vec{A}}{\partial t^2} - \vec{\nabla} \left( \vec{\nabla} \cdot \vec{A} + \frac{1}{c} \frac{\partial \Phi}{\partial t} \right) = -\frac{4\pi}{c} \vec{J}$$

On imposing the Lorenz gauge condition  $\vec{\nabla} \cdot \vec{A} + \frac{1}{c}(\partial\Phi/\partial t) = 0$  one gets the desired equation,

$$\nabla^2 \vec{A} - \frac{1}{c^2} \frac{\partial^2 \vec{A}}{\partial t^2} \vec{\nabla} = -\frac{4\pi}{c} \vec{J}$$

b) We know that  $\left[ \vec{J}(\vec{x}', t') \right]_{ret} = \vec{J}(\vec{x}', t' = t - |\vec{x} - \vec{x}'|/c)$ , so for the given sinusoidal current,

$$\vec{A}(\vec{x}, t) = \frac{e^{-i\omega t}}{c} \int d^3x' \frac{\vec{J}(\vec{x}') e^{ik|\vec{x} - \vec{x}'|}}{|\vec{x} - \vec{x}'|}$$

where  $k = \omega/c = 2\pi/\lambda$ . There are three length scales in the problem: 1) the linear extension of the current distribution denoted by  $d$  (then, with the origin of the coordinate system chosen within the current distribution, one has  $x' \lesssim d$ ), 2) the length  $\lambda$  which is the distance that a signal travels during one oscillation of the source (note that  $2\pi/\omega = T$  is the time period of the oscillating source), 3) the distance to the observer denoted by  $x = |\vec{x}|$ . For a well localized source, we always assume that  $d \ll x, \lambda$ . Now, the “near zone” is characterized by  $d \ll x \ll \lambda$ . We then make the approximation  $k|\vec{x} - \vec{x}'| \sim k|\vec{x}| \ll 1$  or  $e^{ik|\vec{x} - \vec{x}'|} \sim 1$ , so that

$$\vec{A}(\vec{x}, t) = \frac{e^{-i\omega t}}{c} \int d^3x' \frac{\vec{J}(\vec{x}')}{|\vec{x} - \vec{x}'|}$$

Except for the overall time modulation, this has the character of a magnetostatic field. The “far zone” is characterized by  $d \ll \lambda \ll x$ . Then we can make the approximation  $|\vec{x} - \vec{x}'| \sim x - \vec{x} \cdot \vec{x}'/x$  and  $1/|\vec{x} - \vec{x}'| \rightarrow 1/x$ , leading to

$$\vec{A}(\vec{x}, t) = \frac{1}{c} \frac{e^{i(kr - \omega t)}}{x} \int d^3x' \vec{J}(\vec{x}') e^{-ik\vec{x} \cdot \vec{x}'/x}$$

The factor in front of the integral shows that this has the character of an expanding spherical wave.

6. (a) Consider the linear transformation  $\tilde{x}^\mu = L^\mu_\nu x^\nu$ . Find the constraint that the invariance of the space-time interval  $(x^0)^2 - (x^1)^2 - (x^2)^2 - (x^3)^2$  imposes on the matrix  $L$ , showing the steps in your calculation clearly.
- (b) If a 4-vector  $V^\mu$  transforms as a contravariant vector under Lorentz transformations, work out the transformation of  $V_\mu = \eta_{\mu\rho} V^\rho$ .
- (c) Show that the two Maxwell equations with sources are contained in the relativistic expression

$$\partial_\mu F^{\mu\nu} = \frac{4\pi}{c} J^\nu$$

**Solution** (points:(5+4+5))

a) In matrix notation, the space-time interval can be written as  $(x^0)^2 - (x^1)^2 - (x^2)^2 - (x^3)^2 = x^T \eta x$ , where,

$$\eta = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}, \quad x = \begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix}, \quad x^T = (x^0 \quad x^1 \quad x^2 \quad x^3)$$

The linear transformation takes the form  $\tilde{x} = Lx$ , where  $L$  is the  $4 \times 4$  matrix with elements  $L^\mu_\nu$  (with  $\mu$  running over rows and  $\nu$  running over the columns of the matrix). For  $x^T$  the transformation reads  $\tilde{x}^T = x^T L^T$ . The invariance of the interval means that  $x^T \eta x = \tilde{x}^T \eta \tilde{x}$  from which the constraint on  $L$  follows as

$$L^T \eta L = \eta$$

b) In matrix notation the relation between  $V_\mu$  and  $V^\mu$  can be written as  $V_\mu = \eta_{\mu\rho} V^\rho = (\eta V)_\mu$  where the column matrix  $V$  is constructed from the components of the contravariant vector. After the transformation, we have  $\tilde{V}_\mu = (\eta \tilde{V})_\mu$  where,  $\tilde{V} = LV$ . Hence we have,  $\tilde{V}_\mu = (\eta LV)_\mu = (\eta L \eta^{-1} \eta V)_\mu$  or,

$$\tilde{V}_\mu = (\eta L \eta^{-1})_\mu{}^\nu V_\nu$$

This is the transformation of a contravariant vector.

c) We start by writing the relativistic expression separately for  $\nu = 0$  and  $\nu = j$  (where  $j$  is a space index),

$$\partial_i F^{i0} = \frac{4\pi}{c} J^0 \quad \partial_0 F^{0j} + \partial_i F^{ij} = \frac{4\pi}{c} J^j$$

where we have used  $F^{00} = 0$ . Now we note that  $J^0 = c\rho$ ,  $F^{i0} = E^i$ ,  $F^{ij} = \epsilon^{ji}_k B^k$  and  $\partial_i F^{ij} = \epsilon^{ji}_k \partial_i B^k = (\vec{\nabla} \times \vec{B})^j$ . Using these we recover the two Maxwell equations,

$$\vec{\nabla} \cdot \vec{E} = 4\pi\rho, \quad \vec{\nabla} \times \vec{B} - \frac{1}{c} \frac{\partial}{\partial t} \vec{E} = \frac{4\pi}{c} \vec{J}$$