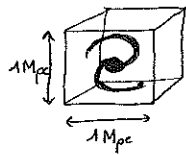


2.1. Suppose that galaxies contain $\sim 10^{11}$ stars and are separated by 1 Mpc
 → Density of the universe?

$$\begin{cases} 1 M_{\odot} \approx 2 \cdot 10^{30} \text{ kg} \\ 1 \text{ pc} \approx 3 \cdot 10^{16} \text{ m} \end{cases}$$



$$\text{Density} = \frac{m}{V}$$

and we have 10^{11} stars / Mpc³

$$\Rightarrow \text{density}_{(e)} = \frac{10^{11} \times M_{\odot}}{(1 \text{ Mpc})^3} = \frac{10^{11} \times 2 \cdot 10^{30} \text{ kg}}{(1 \text{ Mpc})^3} = 2 \cdot 10^{41} \text{ kg / Mpc}^3$$

$$1 \text{ Mpc} = 10^6 \text{ pc} = 3 \cdot 10^{22} \text{ m}$$

$$\begin{aligned} \Rightarrow 1 \text{ Mpc}^3 &= (3 \cdot 10^{22})^3 \text{ m}^3 \\ &= 27 \cdot 10^{66} \text{ m}^3 \\ &\approx 10^{67} \text{ m}^3 \end{aligned}$$

$$\Rightarrow \rho \sim \frac{10^{41} \text{ kg}}{10^{67} \text{ m}^3} = 10^{41-67} \text{ kg} \cdot \text{m}^{-3}$$

$$\Rightarrow \boxed{\rho \sim 10^{-26} \text{ kg} \cdot \text{m}^{-3}} = \text{density of the universe}$$

Density of the Earth: $\rho_{\text{Earth}} = 5,51 \text{ g} \cdot \text{cm}^{-3}$
 $= 5,51 \cdot 10^{-3} \cdot 10^6 \text{ kg} \cdot \text{m}^{-3}$
 $= 5,51 \cdot 10^3 \text{ kg} \cdot \text{m}^{-3}$

So $\boxed{\rho_{\text{Earth}} \gg \rho_{\text{universe}}}$

2.3. If the universe was not electrically neutral (ie. nb protons \neq nb electrons), the electromagnetic force (much stronger than gravitational force) would dominate the structure of the universe.

Evidence suggests that gravity is in fact the dominant force on the large scales.

2.4. Frequency of a photon with energy 13,6 eV?

$$E = h f \Rightarrow f = \frac{E}{h}$$

Constants from the book and units:

$$\begin{cases} h = \frac{h}{2\pi} = 1,055 \cdot 10^{-34} \text{ m}^2 \cdot \text{kg} \cdot \text{s}^{-1} \\ 1 \text{ kg} \cdot \text{m}^2 \cdot \text{s}^{-2} = 1 \text{ J} \\ 1 \text{ eV} = 1,602 \cdot 10^{-19} \text{ J} \end{cases}$$

$$\Rightarrow h = 2\pi \times \frac{1,055 \cdot 10^{-34}}{1,602 \cdot 10^{-19}} \text{ eV} \cdot \text{s}$$

$$\Rightarrow f = \frac{13,6 \times 1,602 \cdot 10^{-19}}{2\pi \times 1,055 \cdot 10^{-34}} \text{ s}^{-1}$$

$$\boxed{f \approx 3,29 \cdot 10^{15} \text{ Hz}}$$

What is the temperature of the photon with mean energy 13,6 eV?

The mean energy of a photon in the black-body distribution is $E_{\text{mean}} \approx 3 k_B T$ (Fig 2.10)

$$k_B = 8,619 \cdot 10^{-5} \text{ eV} \cdot \text{K}^{-1}$$

$$T \approx \frac{13,6}{3 \times 8,619 \cdot 10^{-5}}$$

$$\boxed{T \approx 52600 \text{ K}}$$

2.5. Peak of the energy density distribution of a black-body is at $f_{\text{peak}} \approx \frac{2,8 k_B T}{h}$

$$\Rightarrow \frac{f_{\text{peak}}}{T} = \text{cte}$$

$$\frac{f_{\text{peak}}}{T} = \frac{2,8 k_B}{h}$$



Constants: $h = 2\pi \times 1,055 \cdot 10^{-34} \text{ m}^2 \cdot \text{kg} \cdot \text{s}^{-1}$

$$k_B = 1,381 \cdot 10^{-23} \text{ J} \cdot \text{K}^{-1}$$

$$= 1,381 \cdot 10^{-23} \text{ kg} \cdot \text{m}^2 \cdot \text{s}^{-2} \cdot \text{K}^{-1}$$

$$\Rightarrow \frac{f_{\text{peak}}}{T} = \frac{2,8 \times 1,381 \cdot 10^{-23}}{2\pi \times 1,055 \cdot 10^{-34}} \left(\frac{\text{kg} \cdot \text{m}^2 \cdot \text{s}^{-2} \cdot \text{K}^{-1}}{\text{m}^2 \cdot \text{kg} \cdot \text{s}^{-1}} = \frac{\text{Hz}}{\text{K}} \right)$$

$$\boxed{\frac{f_{\text{peak}}}{T} \approx 5,8 \cdot 10^{10} \text{ K}^{-1} \cdot \text{s}^{-1}} \quad (\text{Hz} \cdot \text{K}^{-1})$$

The Sun: black-body at $T \approx 5800 \text{ K}$

$$\Rightarrow \boxed{f_{\text{peak}} \approx 5800 \times 5,8 \cdot 10^{10} \approx 3,4 \cdot 10^{14} \text{ Hz}}$$

Where in the EM spectrum?

$$\boxed{\lambda = \frac{c}{f} = \frac{299\,792\,458}{3,4 \cdot 10^{14}} \approx 8,82 \cdot 10^{-7} \text{ m} = 882 \text{ nm}}$$

(In the solution the author explains that he wanted it to be in the visible light, which is where the Sun is actually emitting, ie. yellow-green).

near IR
 (750 nm → 100 μm)

2.6. CMB = black-body spectrum at $T = 2,725$ K.

• peak frequency?

$$f_{\text{peak}} \approx \frac{2,8 \text{ k}B T}{h} = \frac{2,8 \times 1,381 \cdot 10^{-23}}{2\pi \times 1,055 \cdot 10^{-34}} \times 2,725$$

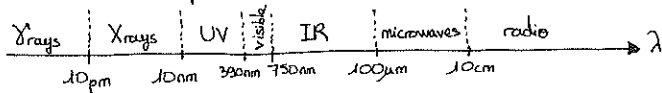
$$f_{\text{peak}} \approx 1,6 \cdot 10^{11} \text{ Hz}$$

• corresponding wavelength?

$$\lambda = \frac{c}{f} = \frac{299\,792\,458}{1,6 \cdot 10^{11}} \Rightarrow \lambda \approx 1,9 \text{ mm}$$

↓
microwaves 😊

Reminder EM spectrum:



• Compare to figure 2.5?

$\lambda \approx 2$ mm means 1 wave every 2 mm

ie. ~ 5 waves per cm

cf. peak of the CMB curve in fig 2.5!

• Total energy density of the CMB?

Given by Stefan-Boltzmann law (cf. lecture):

$$E_{\text{rad}} = \alpha T^4 \quad \text{with } \alpha = \frac{\pi^2 K_B^4}{15 h^3 c^3} = 7,565 \cdot 10^{-16} \text{ J} \cdot \text{m}^{-3} \cdot \text{K}^4$$

$$E_{\text{rad}} = 7,565 \cdot 10^{-16} \times (2,725)^4$$

$$E_{\text{rad}} \approx 4,17 \cdot 10^{-14} \text{ J} \cdot \text{m}^{-3}$$

dimension analysis:

$$[E_{\text{rad}}] = \frac{(\text{kg} \cdot \text{m}^2 \cdot \text{s}^{-2} \cdot \text{K}^{-1})^4}{(\text{m}^2 \cdot \text{kg} \cdot \text{s}^{-1})^3 (\text{m} \cdot \text{s}^{-1})^3} \cdot \text{K}^4$$

$$= \frac{\text{kg}^4 \cdot \text{m}^8 \cdot \text{s}^{-8}}{\text{m}^6 \cdot \text{kg}^3 \cdot \text{s}^{-3} \cdot \text{m}^3 \cdot \text{s}^{-3}}$$

$$= \text{kg} \cdot \text{m}^{-1} \cdot \text{s}^{-2}$$

$$= \text{kg} \cdot \text{m}^2 \cdot \text{s}^{-2} \cdot \text{m}^{-3}$$

$$= \text{J} \cdot \text{m}^{-3}$$